

Hw3, 9.3, 9.5, 9.11, 9.14, 9.16a

9.3

In the near region the potentials are just like the static case with the time dependence $\exp(-i\omega t)$. The potential due to a sphere with opposite voltages is given by equation 3.36 as

$$\Phi = \exp(-i\omega t)V \left(\frac{3a^2}{2r^2} P_1(\cos \theta) + \dots \right)$$

From this expansion we can read off the dipole term

$$\vec{p} = 6\pi\epsilon_0 V \exp(-i\omega t) a^2 \hat{z}$$

Using 9.19 we find the fields in the radiation zone to be

$$H = -\frac{3\epsilon_0 V c a^2 k^2 \sin \theta \exp(ikr - i\omega t)}{2r} \hat{\phi}$$

$$E = -\frac{3V a^2 k^2 \sin \theta \exp(ikr - i\omega t)}{2r} \hat{\theta}$$

The time averaged angular power distribution is

$$\frac{dP}{d\Omega} = \frac{9V^2 a^4}{8Z_0} \sin^2 \theta \cos^2(\omega t),$$

which can be integrated to yield the total power

$$P = \frac{3\pi a^4 k^4 V^2}{Z_0} \cos^2(\omega t)$$

9.5

Keeping only the 1st term in 9.9, we have

$$A = \frac{\mu_0 \exp(ikr)}{4\pi r} \int J(x') d^3x',$$

which can be manipulated as done in the book into

$$A = -\frac{i\mu_0\omega}{4\pi} \frac{\vec{p} \exp(ikr)}{r}.$$

Starting with the top equation on page 410, we can expand $\exp(ik|x - x'|)/|x - x'|$ to get an expression for the scalar potential very much like equation 9.6

$$\Phi = \frac{1}{4\pi\epsilon_0} \sum_{lm} \frac{4\pi}{2l+1} Y_{lm}(\theta, \phi) \frac{\exp(ikr)}{r^{l+1}} (1 + a_1(ikr) + \dots) \int \rho(x') r'^l Y_{lm}^*(\theta', \phi') d^3x',$$

where the omitted terms come from the expansion of $j_l(kr')$. Keeping only the $l = 1$ term, the potential becomes

$$\Phi \approx \frac{1}{4\pi\epsilon_0} \frac{4\pi}{3} \frac{\exp(ikr)}{r^2} (1 - ikr) \sum_m Y_{1m} \int \rho(x') r' Y_{1m}^*(\theta', \phi') d^3x'$$

Using the addition theorem, equation 3.62, the RHS can be transformed into

$$\Phi \approx \frac{1}{4\pi\epsilon_0} \frac{\exp(ikr)}{r^2} (1 - ikr) \hat{n} \cdot \vec{p}$$

From the given expressions for the potentials, we can use

$$H = \frac{1}{\mu_0} \nabla \times A$$

$$E = -\nabla \Phi - \frac{\partial A}{\partial t}$$

to find the fields. With the details omitted, they are simply

$$H = \frac{ck^2 \exp(ikr)}{4\pi r} \left(1 + \frac{i}{kr}\right) n \times p$$

$$E = \frac{k^2 \exp(ikr)}{4\pi\epsilon_0 r} (n \times p) \times n$$

9.11

This charge configuration, as we have learned from last quarter, is a quadrupole with components

$$Q_{11} = Q_{22} = 2qa^2 \cos^2 \omega t$$

$$Q_{33} = -2Q_{11}$$

By equation 9.45, we have

$$\frac{dP}{d\Omega} = \frac{Z_0 c^2 k^6 a^4 q^3}{32\pi^2} \cos^4 \omega t \sin^2 \theta \cos^2 \theta,$$

which can be integrated to yield

$$P = \frac{Z_0 c^2 k^6 q^2 a^4}{60\pi} \cos^4 \omega t$$

9.14

The current density is

$$J = \frac{I_0 \exp(i\omega t)}{2\pi a} \delta(r - a) \delta(\cos \theta) \hat{\phi}$$

In the radiation zone

$$A = \frac{\mu_0 \exp(ikr)}{4\pi r} \int J(x') \exp(-ikn \cdot x') d^3 x'$$

$$= \frac{\mu_0 a \exp(ikr)}{4\pi r} \int_0^{2\pi} \exp(-ika \sin \theta \cos \phi') \cos \phi' d\phi'$$

$$= \frac{\mu_0 a I \exp(ikr)}{2i r} J_1(ak \sin \theta) \hat{\phi}$$

Because $H \sim \nabla \times A$, keeping only terms of $1/r$, we have

$$H = \frac{-Iak \exp(ikr)}{2ir} J_1(ak \sin \theta) \hat{\theta}$$

and

$$E = \frac{-Iak \exp(ikr) Z_0}{2r} J_1(ak \sin \theta) \hat{\phi}$$

Note that the magnetic and electric fields are out of phase and at 90 degrees to each other. It follows

$$\frac{dP}{d\Omega} = \frac{I^2 a^2 k^2 Z_0 J_1^2(ak \sin \theta)}{8}$$

The lowest nonvanishing moment is

$$M_{10} = \frac{Ia^2 \sqrt{3\pi}}{2}$$

9.16

The current density is

$$\vec{J} = \hat{z} I \sin\left(\frac{2\pi z}{d}\right)$$

for $-d/2 \leq z \leq d/2$. By equation 9.8, we have

$$A = \frac{\mu_0 \exp(ikr)}{4\pi r} \int J(x') \exp(-ikn \cdot x') d^3x',$$

which can be simplified to

$$\begin{aligned} A &= \hat{z} \frac{\mu_0 I \exp(ikr)}{4\pi r} \int_{-d/2}^{d/2} \sin(kz) \exp(-ikz' \cos \theta) \\ &= \hat{z} \frac{\mu_0 I d \exp(ikr) \sin(\pi \cos \theta)}{4\pi^2 r \sin^2 \theta} \end{aligned}$$

Since

$$\begin{aligned} H &= ik \hat{n} \times A / \mu_0 \\ E &= ik Z_0 (n \times A) \times n / \mu_0 \end{aligned}$$

the power distribution is

$$\begin{aligned} \frac{dP}{d\Omega} &= \frac{1}{2} \text{Re}[r^2 n \cdot E \times H^*] \\ &= \frac{1}{2} r^2 \frac{k^2 Z_0}{\mu_0^2} |n \times A|^2 \\ &= \frac{Z_0 I^2 \sin^2(\pi \cos \theta)}{8\pi^2 \sin^2 \theta} \end{aligned}$$

We see that there is no power parallel or perpendicular to the antenna. The maxima are at 45 degrees with it.